

# QUADRIVIUM

## ASSIGNMENT 13A:

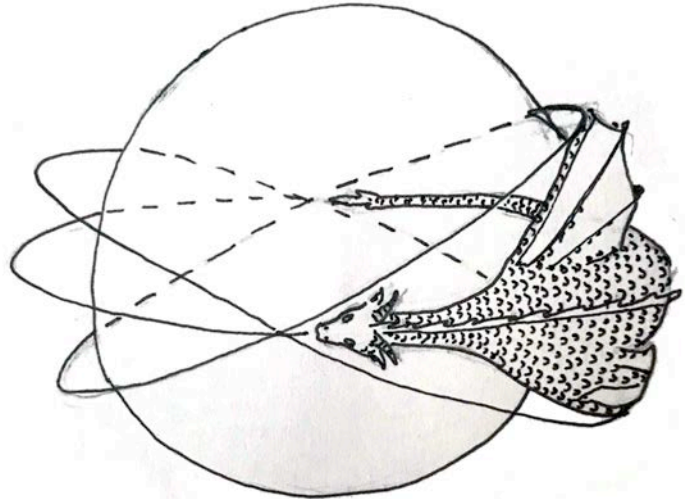
### ASTRO-4

FOR

-ANNANDALE- TUESDAY, APRIL 30<sup>TH</sup>

OR

-FISHKILL- WEDNESDAY, MAY 8<sup>TH</sup>



"...when the moon at full is in the head or tail of the dragon beneath the nadir of the sun, then the earth is interposed between sun and moon, and the cone of the earth's shadow falls on the body of the moon."

-Sacrobosco on the lunar eclipse.  
Izzy's interpretation.

From Grant's Sourcebook (in the Reader, Volume II).

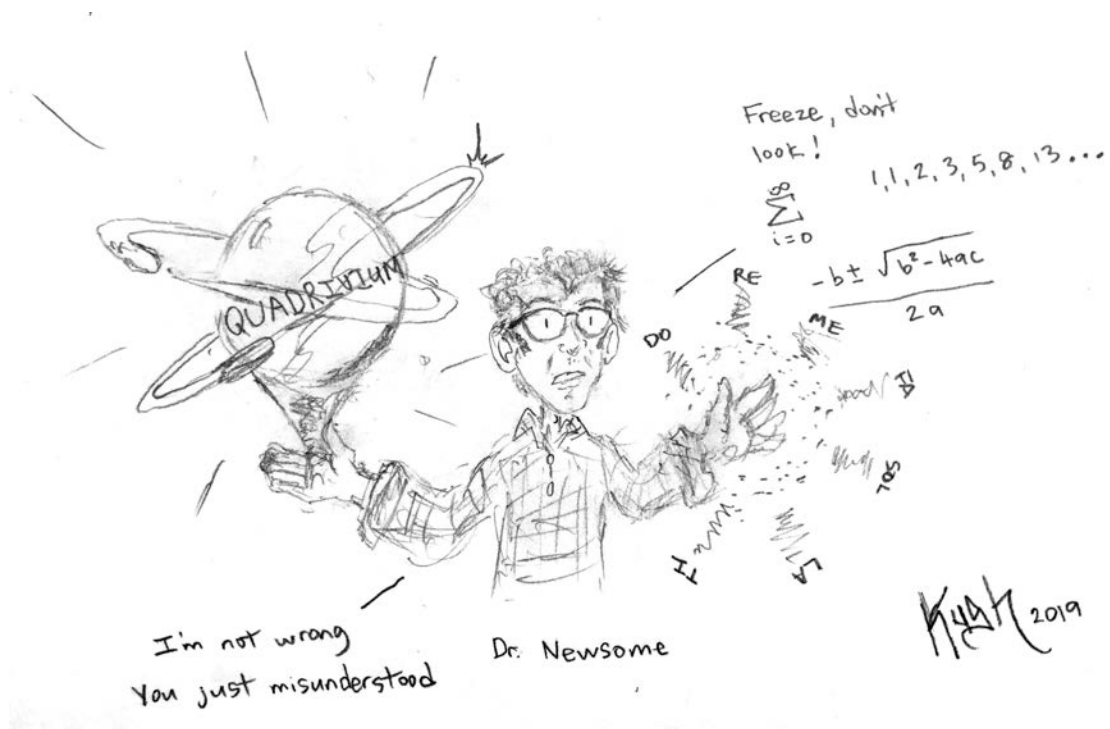
-Read Ch. 66, Oresme's "Attack on Astrology" (pp. 488-494).

-Read Ch. 67, "On the Possible Diurnal Rotation of the Earth" and Ptolemy's "The Immobility of the Earth in the Center of the World" (pp. 494-495).

-Read Ch. 67, Oresme's "The Compatibility of the Earth's Diurnal Rotation..." (pp. 503-510).

-Read Chapters 1 and 2 from Yoko Ogawa's *The Housekeeper and the Professor*. [You bought this at the bookstore (I presume).]

-Do the attached homework.





**HW-13a**  
**Canones de motibus & Tabulae Mediorum Motuum...**  
*- Canon of Motions & Tables of Mean Motions...*  
*by Prosdocimo de Beldemandis (ca. 1370-1428)*  
 folio15r from MS-2284-Bologna- Bib. Universitaria.

**Esidium tabule mediorum motuum planetarum annis christi expansis ab unitate asque ad 28**  
 Likewise, the table of mean motions of the planets for the years of christ expanding from unity [1] to 28.

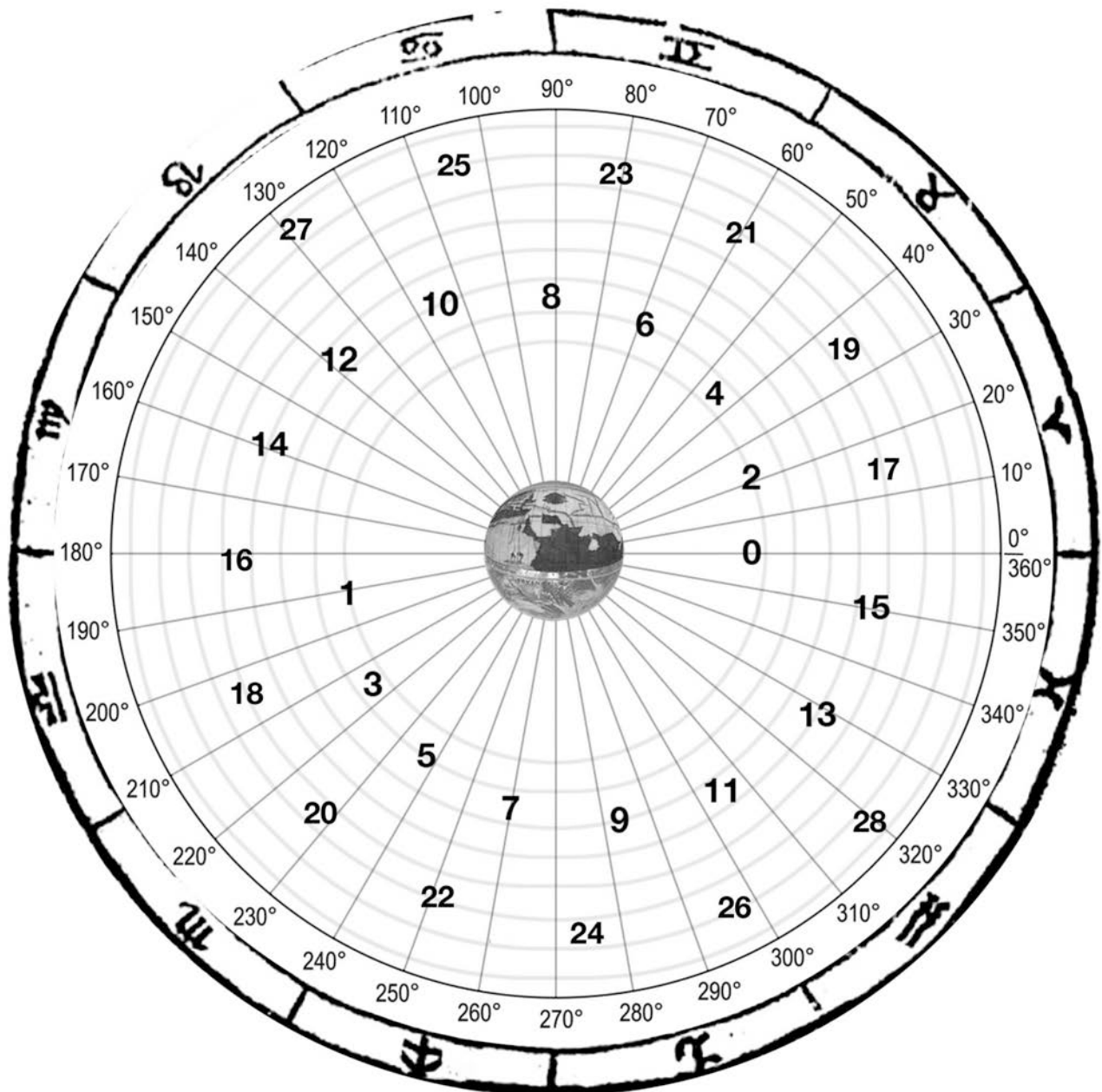
Medio motu Martis Mean Motion of Mars				Modern Analysis		Year-to-Year Difference
sign = 60°	1°	1/60°	1/3600°	Year	Conversion to degrees in decimal notation	
sig.	Grad.	min.	sec.			
o	o	o	o	0	0.0000	
3	11	17	5	1	191.2847	Year <sub>1</sub> - Year <sub>0</sub> = 191.2847
o	22	34	10	2	22.5694	Year <sub>2</sub> - Year <sub>1</sub> = 191.2847
3	33	51	16	3	213.8544	Year <sub>3</sub> - Year <sub>2</sub> = 191.2850
o	45	39	48	4	45.6633	Year <sub>4</sub> - Year <sub>3</sub> = 191.8089
3	56	56	53	5	236.9481	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2848
1	8	13	58	6	68.2328	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2847
4	19	31	3	7	259.5175	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2847
1	31	19	35	8	91.3264	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.8089
4	42	36	40	9	282.6111	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2847
1	53	53	46	10	113.8961	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2850
5	5	10	51	11	305.1808	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2847
2	16	59	23	12	136.9897	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.8089
5	28	16	27	13	328.2742	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2844
2	39	33	33	14	159.5592	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2850
5	50	50	38	15	350.8439	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2847
3	2	39	10	16	182.6528	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.8089
o	13	56	15	17	13.9375	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2847
3	25	13	21	18	205.2225	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2850
o	36	30	26	19	36.5072	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2847
3	48	18	58	20	228.3161	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.8089
o	59	36	3	21	59.6008	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2847
4	10	53	8	22	250.8856	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2847
1	22	10	14	23	82.1706	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2850
4	33	58	45	24	273.9792	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.8086
1	45	15	51	25	105.2642	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2850
4	56	32	56	26	296.5489	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2847
2	7	50	1	27	127.8336	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.2847
5	19	38	33	28	319.6425	Year <sub>n</sub> - Year <sub>n-1</sub> = 191.8089
					<b>Average:</b>	<b>191.4207</b>

Here we see the median motion of Mars over 28 years. This is the general motion of Mars on its deferent. Recall that the deferent is the circle that carries the epicycles of Mars. It "ferries" all the Martian geometrical apparatus. The modern analysis converts all the sexagesimal numbers to modern decimal degrees, finds the angular motion per year, and then averages this angular motion.

Mars' average orbital velocity on the deferent is ca. 191.42° per year. Using the Rule of 3 you can figure out how many years for a full orbit, its orbital period.

$$\frac{191.42^\circ}{1 \text{ year}} = \frac{360^\circ}{x} \rightarrow x \cong 1.881 \text{ years} \rightarrow 1.881(365.25) \cong 687 \text{ days.}$$

The modern measurement for the period for Mars' orbit around the sun is... you guessed it... 687 days. Medieval astronomy was pretty accurate.



In this chart the earth is in the center and the location of Mars is indicated by year: 1<sup>st</sup> year, 2<sup>nd</sup> year, 3<sup>rd</sup> year, etc.

This is where Mars will be within the zodiac on a yearly basis over a span of 28 years.

The radial distance of Mars is not accurately indicated in this chart.

The numbers spiral out so that their angular pattern can be seen and so that the numbers don't have to be written on top of each other.

### 13a: Homework Assignment

Figure out the orbital period of Venus in years and then in days from the following data.

DATA FROM ALFONSINE TABLE						
Sexagesimal						
A	B	C	D	E	F	G
$60^2$ 3600°	$60^1$ 60° sig.	$60^0$ 1° Grad.	Year	Angular rotation in Degrees (in decimal)	Difference in Degrees (decimal) $degrees_{year\ n} - degrees_{n-1}$	Location on Diagram
0	3	45	I	225°	n.a.	225°
0	13	30	II			
0	23	15	III			
0	33	0	IV	1980°		
0	42	45	V	2565°	585°	45°
0	52	30	VI			
1	02	15	VII			
1	12	1	VIII		586°	
1	21	46	IX			
1	31	31	X	5491°		91°
1	41	16	XI			
1	51	2	XII			
					Average:	

#### Instructions

- Finish Column E: Convert the Sexagesimal angles into decimal angles.  
E.g. In year 5 →  $42:45 = 42(60^\circ) + 45(1^\circ) = 2565^\circ$ . Put your results in column E.  
Note: I converted the original manuscript data in Columns A, B, and C into rotation angles.  
For example a resultant decimal angle measuring  $2565^\circ$  is 7 complete rotations plus  $45^\circ$ ,  $7 \times 360^\circ + 45^\circ$ .
- Finish Column F: Figure out the angular difference from one year to the next by subtracting each year's angular rotation from the previous year.  
E.g. Year 5 differs from year 4 →  $2565^\circ - 1980^\circ = 585^\circ$ . Put these results in column F.
- Finish Column G and plot points: Locate Venus on the diagram below. Here's how....For year 5, the angle of rotation [column E] is  $2565^\circ$ . To figure out where that is on a  $360^\circ$  diagram, just divide  $2565^\circ$  by 360. The result is 7.125. Venus completed 7 rotations and 0.125 more. The location on the diagram is just the 0.125 part. 0.125 times  $360^\circ$  is the location.

$$0.125(360^\circ) = 45^\circ.$$

Then plot these points with the year-number on the diagram below. I plotted a few to show you how.

Another way to find the location is to take the angle of rotation from column D and subtract 360° over and over again until you get a number below 360°.

$$\text{E.g. } 2565^\circ - 360 - 360 - 360 - 360 - 360 - 360 - 360 = 45^\circ.$$

or

$$2565^\circ - 7(360^\circ) = 45^\circ$$

4) Determine the length of time it takes Venus to complete one orbit. First find the average of all the differences of column F. This is the average angular rotation of Venus per year. Use the Rule of 3 to figure out how long it takes Venus to complete one (360°) rotation.

$$\frac{\text{Column F Average}}{1 \text{ year}} = \frac{360^\circ}{x} \rightarrow x \cong \text{ years} \rightarrow x(365.25) \cong \text{ days.}$$

Compare this orbital period with the modern value for the orbital period of Venus. Look it up.

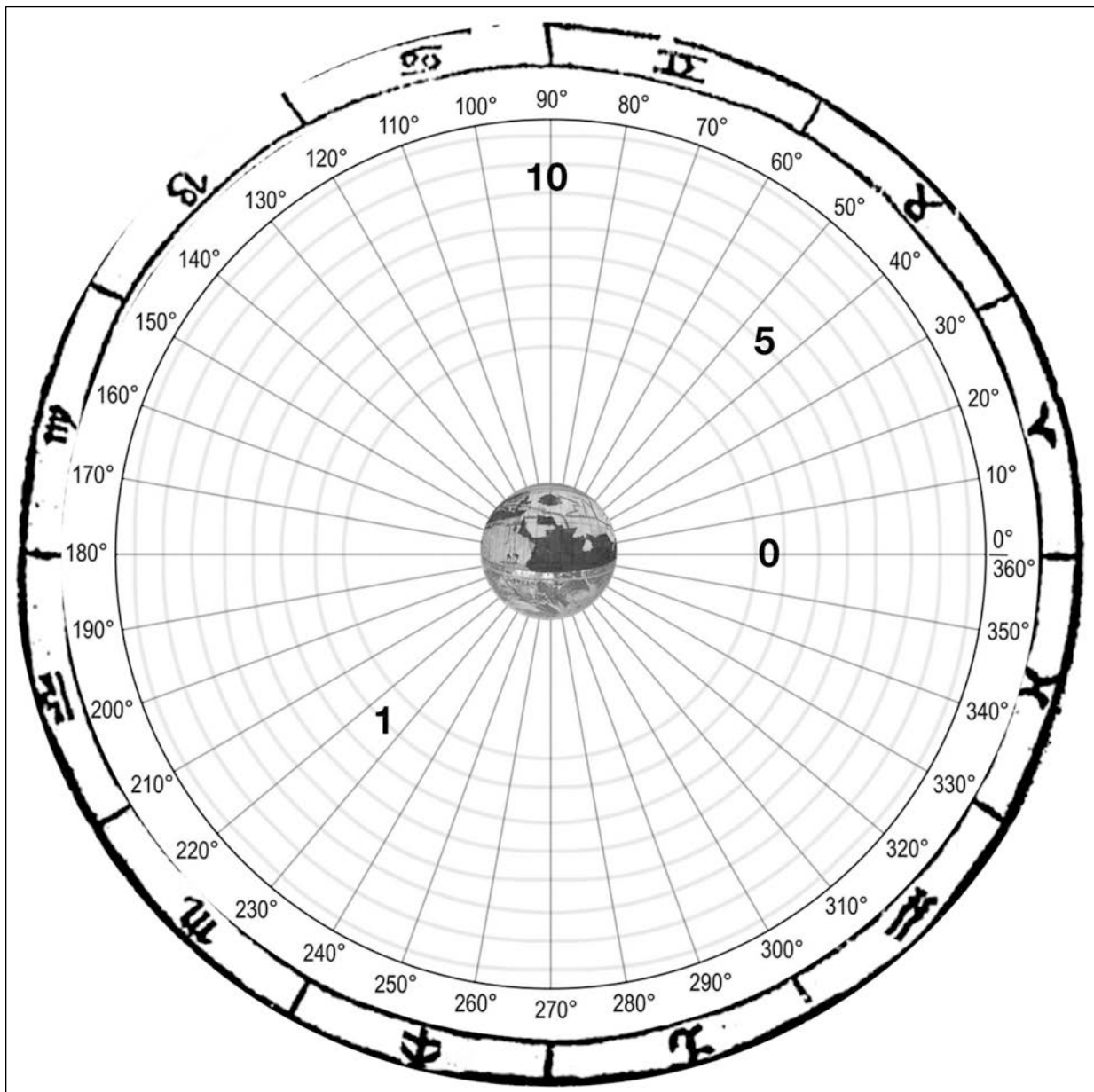


Diagram of where Venus will be seen within the zodiac every year over a span of 12 years.